

Orange Flame from Active Nitrogen and Oxygen in the Absence of a Metal Catalyst Resulting from Collisional Intersystem Crossing into $N_2(B^3\Pi_g)$

Efstathios Kamaratos

Physical Chemistry Sector, University of Ioannina, GR 45110 Ioannina, Greece, and LASIM,[†] Bât. 205, Université Lyon I & CNRS, 43 Bd. du 11 Nov. 1918, F 69620 Villeurbanne Cedex, France

Received: July 12, 1996; In Final Form: November 14, 1996[⊗]

The discovery of an orange flame is reported that appears also at higher pressures of active nitrogen and oxygen without the presence of a metal catalyst unlike previous research in over 30 years. Enhancement of emissions of the first positive band system of nitrogen was observed from various vibrational levels of the $N_2(B^3\Pi_g)$ state. Arguments to identify the species responsible and implications to previous research of this phenomenon of homogeneous excitation and collisional intersystem crossing into $N_2(B^3\Pi_g)$ induced by excited O_2 are presented. Its probing for lasing is suggested.

Introduction

For more than 30 years extensive research work has been carried out on radiationless energy transfer and emissions of radiation attributed to metal surface catalyzed (heterogeneous, in general) formation of species. Such species of great importance also to various applications are excited molecular states of nitrogen,^{1–4} oxygen,^{5–8} nitric oxide,^{2,3,8,9} and nitrogen dioxide.^{5,8,10} They are observed in a gas flow of active nitrogen (or nitric oxide to form nitrogen dioxide) and active oxygen premixed or mixed after activation or separately (e.g., as in the study of excited molecular oxygen states) in laboratories on Earth or in space (e.g., the shuttle glow).¹¹ Also innumerable investigations of the nitrogen afterglow¹² with and without oxygen species have been carried out for more than 100 years for basic research and for a broad spectrum of applications such as atmospheric processes,^{12,13} discharges and plasmas,¹⁴ laser excitation,¹⁵ and semiconductor processing.¹⁶ In some of these investigations in the presence of a metal, enhancement of background nitrogen afterglow emissions in the form of a red flame was reported,^{1–4} when a mixed stream of nitrogen atoms and discharged oxygen was passed over a metal surface. It was initially¹ ascribed to metal-catalyzed formation of $NO(B^2\Pi)$ from N and O, which upon reacting with N atoms led to red chemiluminescence. However in later work the red flame appeared even when a metal was placed in the active nitrogen flow upstream of the discharged oxygen flow.¹⁷ The observed enhancement of emissions from $v = 6, 7,$ and 8 of $N_2(B^3\Pi_g)$ was ascribed¹⁷ to probably excited O_2 , more likely $O_2(a^1\Delta_g)$, interacting in the gas phase with $N_2(A^3\Sigma_u^+)$ formed at the metal surface. The red flame could not be observed unless both a metal and discharged oxygen were added to the active nitrogen flow at <1 Torr. Further research on the mechanism of this surface-catalyzed flame was done more recently.¹⁸

Experimental Section

In the present work the appearance of a bright orange flame is reported from the mixing of active nitrogen and active oxygen gas streams also at much higher N_2 and O_2 pressures and without the presence of a metal for the first time. The flame predominated over the background afterglow emissions and could also be observed with the lab lights on. The flame extended axially

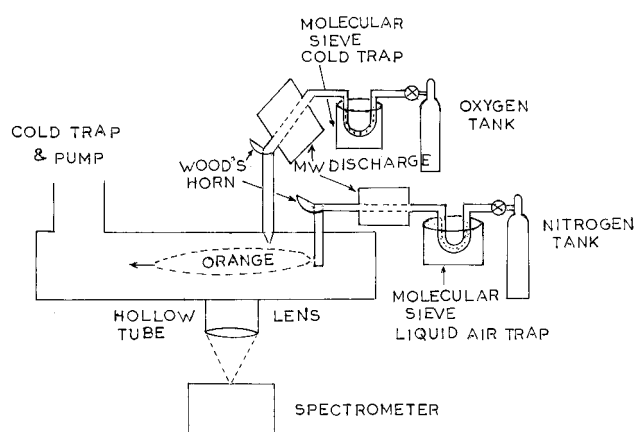


Figure 1. Schematic of the apparatus.

in the reactor and had zero intensity near the walls, thus indicating no contribution from heterogeneous wall effects. The active species were generated by flowing nitrogen and oxygen, separately, through 2450 MHz microwave (mw) discharges. The flows were mixed inside a cylindrical flow reactor initially made of Pyrex^{19a,b} and later of quartz.^{19c} Figure 1 shows a schematic of the apparatus. Gas flows in the range of about 3–10 bar cm^3/s were used. The reactors had an i.d. of 2.0–3.8 cm. The afterglow emissions were observed in variable low resolution (to 3 Å) with spectrometers, filters, and detectors (Ge, Si, and photomultipliers with S-20, S-5, and 605s photocathodes) from IR to UV. Scattered light to the detectors used from the mw discharges used and external sources was eliminated with standard techniques (e.g., by using Wood's horns, black paint, black cloth, etc.). Gratings blazed in the first order at 600 nm and at 1 μm were used. A requirement for the appearance of the orange flame with band intensity I over the background emissions in the nitrogen afterglow with band intensity I_0 was the activation of the oxygen species in the oxygen flow entering the reactor. When the oxygen mw discharge was off, the orange flame disappeared. Only the background emissions remained from discharged N_2 mixed with undischarged O_2 with band intensity I_0 ($\Delta I = I - I_0$).

Results and Discussion

A spectroscopic investigation has been conducted of excitation and deexcitation of molecular species in gas mixtures of

[†] Address during a sabbatical leave.

[⊗] Abstract published in *Advance ACS Abstracts*, January 15, 1997.

active nitrogen and active oxygen. The investigation concentrated on the first positive band system (1pbs) of nitrogen, $N_2(B^3\Pi_g) \rightarrow N_2(A^3\Sigma_u^+)$, although other band systems (O_2 , NO, N_2) were preliminarily examined and further work on them is planned. The extent of the enhancement of emissions depended on the proximity to the mixing region of the active species and the N_2 and O_2 pressure, which was sometimes >20 times higher than that in the past.¹⁷ The enhancement was observed in the flame and downstream from it (by tens of milliseconds). A plot of the relative enhancement $\Delta I/I_0$ of the intensity I_0 of radiation from the background 1pbs emissions in the presence of undischarged O_2 , which is enhanced by active oxygen to intensity I , vs $N_2(B, \nu)$ is shown for some cases for observation normal to the flow direction: downstream of the mixing region (over 50 ms in flow time) at 5.5 mbar nitrogen and 0.1 mbar oxygen in Figure 2a; and at higher pressures (15 mbar N_2 ; 0.9 mbar O_2 , at a flow velocity of about 0.2 m/s) closer (<10 ms) to the mixing region, where the flame originated, in Figure 2b. Enhancement from many vibrational levels up to $\nu = 18$ is shown in Figure 2b for the case of a nitrogen "pink" afterglow. Figure 2a shows no emission from $\nu \geq 10$ in the present homogeneous gaseous excitation as in the heterogeneous excitation of N_2 species,¹⁷ where emissions from $\nu \geq 9$ were reported suppressed. This may indicate similar excitation of the precursor species that populate $N_2(B, \nu)$ in the nitrogen afterglow under the action of active oxygen in both cases. Much higher enhancements than those in Figure 2b were also observed, e.g., even up to ≥ 50 times higher for emissions from $N_2(B, \nu=6)$. The much larger relative enhancement from $\nu = 6$ was evident in all cases as in Figure 2. At distances downstream from the discharges (≈ 0.1 s flow time) and at total pressures up to 10.7 mbar with $1 < N_2/O_2 < 100$ a similar preferential enhancement from $\nu = 6$ was observed. In some experiments when the O_2 pressure was increased with the N_2 pressure kept constant, the relative increase in the emissions from $\nu \leq 6$ was larger than that from $\nu = 11$. Spectra of the 1pbs of N_2 are shown in Figures 3 and 4.

Experiments to obtain information on the oxygen and nitrogen species responsible for the 1pbs enhancement included monitoring the concentrations of $O_2(a^1\Delta_g)$ and $O_2(b^1\Sigma_g^+)$ and that of O atoms. The $\nu = 0 \rightarrow \nu' = 0$ emissions to $O_2(X^3\Sigma_g^-)$ at 1268 and 762 nm and the NO_2^* emission in the presence of $NO^{20,21}$ were measured as a function of the oxygen discharge mw power, before and after mixing with nitrogen. These were compared with the corresponding change of the 1pbs enhancement. The $O_2(^1\Sigma)$ signal was very small and could be observed photometrically only without N_2 . In a N_2 Lewis–Rayleigh afterglow the 1pbs emission intensity is proportional¹² to $[N]^2$. Therefore, since the observed 1pbs enhancement is nonuniform (Figure 2) even under the normal afterglow conditions (emissions only from $\nu \leq 12$), it is deduced that it is not the usual N atom recombination that is mainly responsible for this phenomenon. NO β emissions were observed in the mixing region, when oxygen was added upstream of the nitrogen inlet. They disappeared in the mixing region and very close to the oxygen inlet, when oxygen was added downstream of nitrogen, even though 1pbs enhancement was observed there. These results suggest that it is some N_2 species that are responsible for fast collisional intersystem energy transfer to $N_2(B)$ by colliding with active oxygen rather than with excited NO, before N atoms can react with O atoms or recombine by O atoms. Thus from these observations of nonuniform, fast 1pbs enhancement from low/high ν of $N_2(B)$, even after times longer than 0.1 s from the discharges, it is surmised that 1pbs enhancement appears by a mechanism of collisional crossing into $N_2(B)$

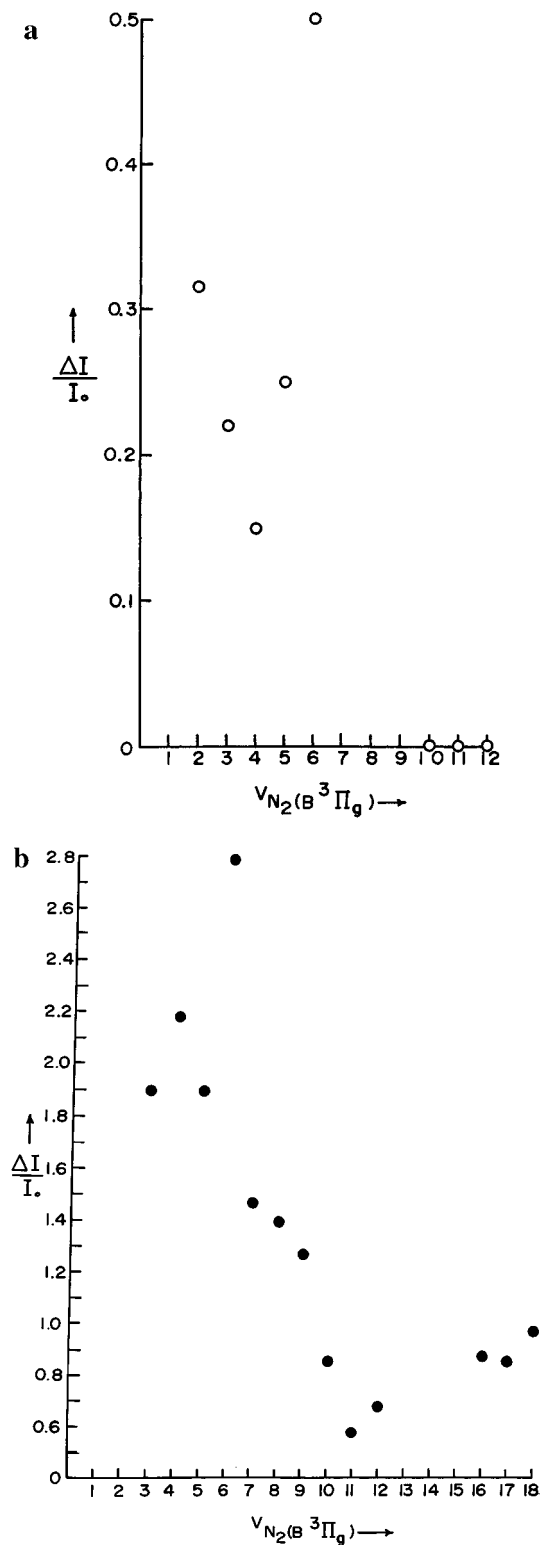


Figure 2. Relative enhancement of first positive band system (1pbs) emissions of N_2 from a gas mixture of discharged nitrogen and separately discharged oxygen in a flow reactor. I represents N_2 1pbs band intensity. I_0 represents N_2 1pbs band intensity, however, from a gas mixture of discharged nitrogen as above and undischarged oxygen in the same flow reactor. $\Delta I = I - I_0$. (a) Downstream of the mixing region. $P_{N_2} = 5.5$ mbar. $P_{O_2} = 0.1$ mbar. (b) Near the mixing region of a "pink" nitrogen afterglow (extending visually downstream almost to the oxygen inlet) with oxygen. $P_{N_2} = 15.0$ mbar. $P_{O_2} = 0.9$ mbar. The maximum absolute deviation of intensities was 10%.

involving collisions of N_2 species with oxygen species. Energetically possible N_2 species are vibrationally excited N_2 , N_2^+ , metastable $N_2(A)$, and other N_2^* species that can be generated

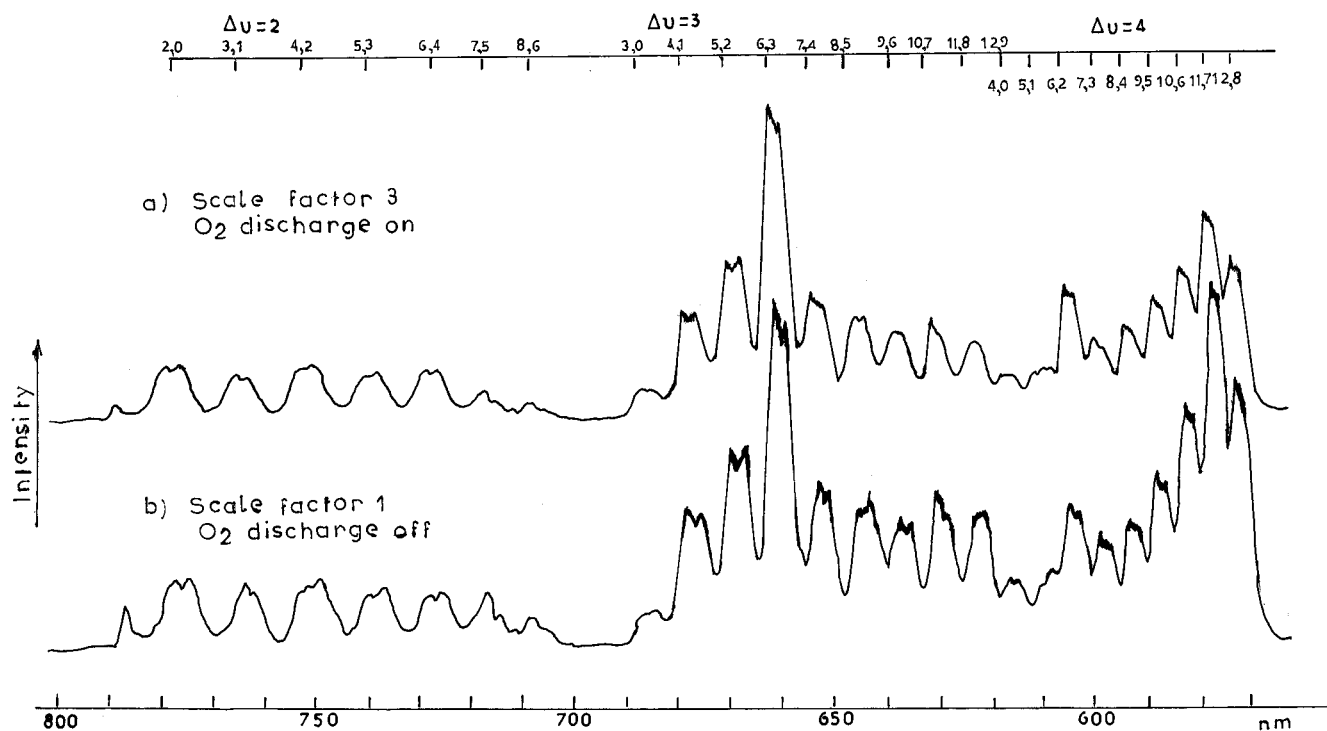


Figure 3. Emission spectra showing the $\Delta\nu = 2$, $\Delta\nu = 3$, and $\Delta\nu = 4$ band sequences of the first positive band system (1pbs) of $\text{N}_2(\text{B } ^3\Pi_g \rightarrow \text{A } ^3\Sigma_u^+)$. The vibrational levels of each band in a 1pbs sequence are shown over each band with the $\text{N}_2(\text{B } ^3\Pi_g)$ vibrational level of each band shown first. The spectra were obtained by adding (a, top spectrum) discharged oxygen to a "pink" nitrogen afterglow with a scale factor of 3 and (b, lower spectrum) undischarged oxygen to the same "pink" nitrogen afterglow with a scale factor of 1. In both cases $P_{\text{N}_2} = 12.0$ mbar and $P_{\text{O}_2} = 0.7$ mbar. The spectra are not corrected for grating and detector nonuniform spectral response.

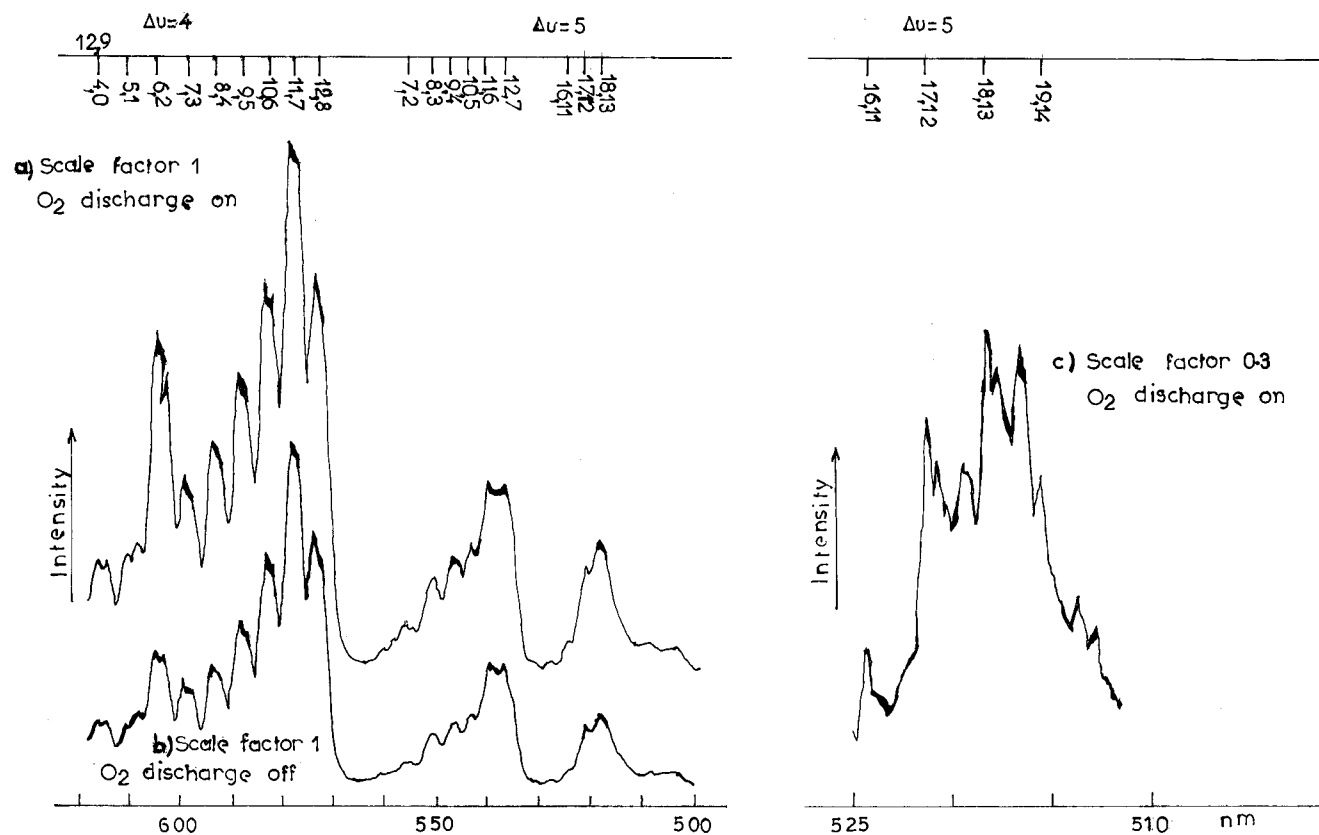


Figure 4. Emission spectra showing the $\Delta\nu = 4$ and $\Delta\nu = 5$ band sequences of N_2 1pbs with (a, top spectrum) and without (b, lower spectrum) discharged oxygen with a scale factor of 1 and the rest of the experimental conditions the same as those in Figure 3. On the right side (c) the emission spectrum with discharged oxygen is shown for some high- ν $\Delta\nu = 5$ bands at higher resolution with a scale factor of 0.3. $P_{\text{N}_2} = 10.3$ mbar and $P_{\text{O}_2} = 0.3$ mbar.

in the N_2 afterglow and may cross into the $\text{N}_2(\text{B})$ state or are coupled to $\text{N}_2(\text{B})$ like $\text{N}_2(\text{W } ^3\Delta_u)$, $\text{N}_2(\text{B}' ^3\Sigma_u^-)$ and $\text{N}_2(\text{A}' ^5\Sigma_g^+)$.¹²

The intensity of emission from a large flame appeared to optimize by the eye at low mw power of the oxygen discharge

and high mw power of the nitrogen discharge. The latter observation was confirmed by quantitative measurements on the intensity of various 1pbs bands. The comparison of the quantitative effect of the change of the oxygen discharge mw power on the excited O₂ concentration and on the 1pbs intensity led to the following results: it showed that the initial concentration of O₂(¹Δ) and O₂(¹Σ) changed similarly to the observed 1pbs intensity, in contrast to [O], which changed rather proportionally to NO emissions (with the flame on). In an experiment discharged oxygen was passed through a piece of quartz tubing with a ring of mercuric oxide film (for O atom removal) in it and subsequently through a ≈5 cm piece of Tygon tubing, before entering the reactor. No change in the orange flame was observed. Spectral scans with and without the mercuric oxide showed no significant changes in visible 1pbs emissions. That O atoms are unlikely to cause this 1pbs enhancement results also from the observations that the flame intensity appeared to optimize at low O₂ discharge mw power (i.e. at low O concentration), as mentioned earlier, and that the O₂ addition to the N₂ afterglow does not change the N₂(B) distribution significantly. However in such a gas mixture there are already O atoms, as expected from the chemical reactions of O₂ with species in the N₂ afterglow and as evidenced by NO* emissions, but no 1pbs enhancement has been observed. O₂[†] (vibrationally excited) is considered unlikely since it may be excited in active N₂ without leading to 1pbs enhancement. Also, when active oxygen flowed through glass wool (for O₂ vibrational relaxation), there was no effect on the flame. It is concluded that electronically excited O₂, O₂^{*}, is the likely candidate for this phenomenon with O₂(¹Δ) the most likely one, because of the very low concentration and the faster quenching rate of O₂(¹Σ) (or of another O₂^{*} state).

About 25 years ago N₂(A) was considered the main N₂ state involved in populating N₂(B) in the nitrogen afterglow in general¹² and in its surface-catalyzed excitation, in particular.¹⁷ However, after the discovery of the optical transitions and the indicated collisional coupling of N₂(B↔W)^{12,13} and during the gradual extension of the available knowledge on the chemical kinetics and dynamics of the quenching of metastable N₂(A) by atoms such as N(⁴S) and O(³P) and molecules such as O₂,^{12,21–24} there has been a shift away from N₂(A).²² This shift turned in favor of N₂(W)¹⁸ as the main state responsible for collisional intersystem crossing into N₂(B). This occurred not only for N₂ metal-catalyzed excitation¹⁸ but also for a low-altitude aurora¹³ and relaxation with coupling of N₂(B).^{25–27} This shift may have led to the indicated unawareness about the possible fast N₂(A) + O₂^{*} interaction that may interfere in the determination of the rate coefficient of the reaction N₂(A) + O,^{21,23–24} although such an interaction had already been proposed for a fast energy transfer process induced by O₂(¹Δ) that populates N₂(B).¹⁷

However experiments have shown two components in N₂(B) fluorescence decay.^{25–28} Also recently the results of molecular beam one-collision crossing into N₂(B,*v*)²⁸ showed (from the different 1pbs spectral distributions at two beam flight-times) that there are at least two species that populate N₂(B) by collisional intersystem crossing: one (or more) of shorter and another one (or more) of longer lifetime. They were ascribed to N₂(W) and N₂(A) on the basis of their shorter and longer lifetime, respectively. The experimental results in Figure 4 of Ottinger and co-workers²⁸ show their 1pbs Δ*v* = 3 sequence in the N₂ afterglow spectrum of an active nitrogen molecular beam, before undergoing any collision. The sequence is almost identical (except for some shift in their wavelength for each band, presumably due to inadvertent error) to that of an aurora

spectrum.¹³ This would be expected of similarly (electron impact) excited N₂ at similar fluorescence times, before long fluorescence or any collisions change it. However after one collision in passing through O₂ or NO or N₂, etc., the molecules in the beam are reported to show (no spectrum for O₂ is shown) large enhancement of their emission with the largest one from N₂(B,*v*=6), as in this work. The spectral distribution after one collision with N₂ also shows that the longer lifetime species ascribed to N₂(A) cross into N₂(B,*v*) for all *v* ≤ 12; however, they cross faster for *v* = 6 and 4. These levels have a small energy gap with the N₂(A,*v*'=16) and N₂(A,*v*'=13). The shorter lifetime species ascribed to N₂(W) appear in the spectrum to cross into N₂(B,*v*) significantly at least for *v* = 5, 6, and 7. These N₂(B) levels are close to the N₂(W,*v*'') ones for *v*' = 6, 7, and 8. Therefore the N₂(A) state cannot be excluded; both N₂(A) and N₂(W) are probably involved in this excitation of N₂(B). Nevertheless, involvement also of some other species cannot be excluded. Such species are N₂[†], through its reported reaction with N₂(A)^{29a} and other N₂^{*} states, as mentioned previously, especially for excitation of N₂(B,*v*>12) induced by O₂^{*}. Collisional intersystem crossing from N₂(B) into N₂(B')^{27b} and from N₂(a ¹Π_g) into (a' ¹Σ_u⁻)²⁹ as well as the absorption^{29c} and laser-induced fluorescence^{29d} after earlier failure^{29c} for N₂(C'' ⁵Π_u) – (A' ⁵Σ_g⁺) were directly observed recently. The previous arguments for this collisional energy transfer lead to the next possible equation for the mechanism O₂^{*} + N₂^{*} → N₂(B,*v*) + O₂' where O₂' may be ground state O₂ or excited O₂; N₂^{*} or N₂[†] may act as a third body.

Interference of this excitation may more or less have affected previous research. For example, all values of rate constants for the reaction of O atoms (from discharged oxygen) with N₂(A) (from discharged nitrogen, or from excitation transfer to N₂ by Ar metastables) may be affected, if indeed N₂(A) interacts with O₂^{*} in this phenomenon, because O₂^{*} is formed along with O atoms in (and outside) the discharge. This should be taken into account also in (upper) atmospheric models²³ or anywhere such N₂(A) interactions matter. Such reactions of O atoms produced even from the reaction NO + N may be affected by this collisional excitation due to potential O₂(¹Δ) formation in heterogeneous recombination of O atoms.³⁰ Speculative estimates of possible error up to 30% from the interaction of N₂(A) with O₂(¹Δ) in some cases were reported^{21,23,24} for the rate constant of the reaction N₂(A) + O. The error may be even higher. It depends on how fast, compared to O₂, O₂^{*} may interact with N₂(A). All values reported for the rate constants for the chemical reaction N + O₂(¹Δ) studied^{31–33} in N₂ afterglows may be at most upper limits, if this reaction occurs at a significant rate, because of the possibility of interference of the energy transfer between O₂(¹Δ) and N₂(A) or some other N₂^{*} species. This energy transfer has also probably caused the too large increase in background emissions and affected the results reported^{24,34} in the study of the reaction of N₂(A) with O from discharged oxygen.

Further Discussion and Conclusion

A flame and/or a simple enhancement of N₂ afterglow emissions caused by active oxygen were also reported in the past in other experiments.^{1–4,17,18,35–37} They show aspects of the same basic phenomenon reported here, i.e., enhancement of the 1pbs due to collisional intersystem excitation transfer into N₂(B) induced by collisions with apparently excited O₂. However, in all of the previous works the occurrence of the 1pbs enhancement was related to some different factor, namely, either to the presence of a metal at <1 Torr, or only to the mixing region of the discharged O₂ and N₂ streams during some of the

TABLE 1: Summary of Research Work That May Have Been Affected by This Excitation to $N_2(B^3\Pi_g)$

| | reference |
|--|-------------------|
| 1. rate of interaction between $N_2(A^3\Sigma_u^+)$ and O | 21, 23, 24, 34 |
| 2. rate of interaction between N and $O_2(a^1\Delta_g)$ | 31–33 |
| 3. exclusion of the role of $N_2(A^3\Sigma_u^+)$ in populating $N_2(B^3\Pi_g)$ | 13, 18, 22, 25–27 |
| 4. mechanism of the $N_2(B^3\Pi_g)$ vibrational redistribution | 35 |

photometric attempts to determine the rate constant of the $O_2(^1\Delta)$ reaction with N atoms in the N_2 afterglow,^{36,37} or to the effect of O (and H) atoms³⁵ in the study of the vibrational redistribution of $N_2(B)$. This was reported to occur only at a N_2 pressure of 0.01–1.0 Torr and was considered to be in agreement with earlier work.³ However, it is contrary to the work later revised¹⁷ and to results in this work. For this reason or for some other unknown reason(s) this homogeneous intersystem crossing into $N_2(B)$ apparently induced by collisions between N_2^* and excited O_2 , as evidenced by nonuniform enhancement of N_2 1pbs emissions in a broad range of total pressure and N_2/O_2 ratios, remained unnoticed. It has probably affected several investigations³⁸ summarized in Table 1, as a recent reinterpretation³⁹ of experimental results^{21–24,31–35} also shows, and it should be taken into account in all pertinent applications. This reinterpretation³⁹ provides additional support for the role of $N_2(A)$ and O_2^* in this excitation. It might even help shed more light on, if not elucidate, the mechanism of the nitrogen afterglow in general and that of the $N(^4S)$ recombination in particular, which are not yet fully resolved,^{12,29} for example, by determining the populating and depopulating effect of O_2^* on various N_2^* states. It was previously conjectured that by variation of the N_2 and O_2 pressure, similar precursor(s) may have been formed in the homogeneous and the heterogeneous N_2 excitation. Much higher N_2 pressures were also used in these experiments compared to <1 Torr used in the past.¹⁷ Through these two factors, i.e., pressure and variability of composition, it may be possible to find conditions for partial or total population inversion and high intensities for a visible laser by employing this phenomenon, i.e., this homogeneous collisional intersystem excitation transfer into $N_2-(B,\nu)$ induced by O_2^* . In fact it was reported that low ν $N_2-(A,\nu)$ states predominate²⁴ over higher ones (e.g., $\nu = 0, 1$ over $\nu = 3$) at high pressure and react with O_2 (and O) slower than higher ν states.^{21,24,34} Near IR lasing of pure N_2 1pbs and of other bands has been reported.¹⁵ There is no visible N_2 1pbs laser or chemical laser.⁴⁰

Acknowledgment. I thank Prof. R. Bacis for the hospitality and the use of his laboratory facilities at the University of Lyon I, France. Partial support is acknowledged from (α) grants to

the University of Ioannina by the General Secretariat for Research and Technology (works 87E Δ 128, 91E Δ 32); (β) grants (works 135, 182, and 435) from the University of Ioannina Research Committee; and (γ) grants to York University, Canada (Profs. H. I. Schiff and C. H. Dugan).

References and Notes

- (1) Reeves, R. R.; Mannella, G. G.; Harteck, P. *J. Chem. Phys.* **1960**, *32*, 946.
- (2) Mannella, G. G.; et al. *J. Chem. Phys.* **1960**, *33*, 636.
- (3) Harteck, P.; et al. *Can. J. Chem.* **1960**, *38*, 1648.
- (4) Mannella, G. G. *Chem. Rev.* **1963**, *63*, 1.
- (5) Harteck, P.; Reeves, R. R. *Discuss. Faraday Soc.* **1964**, *37*, 82.
- (6) Ogryzlo, E. A. *J. Photochem.* **1984**, *25*, 389.
- (7) Sharpless, R. L.; et al. *J. Chem. Phys.* **1989**, *91*, 7936.
- (8) Kenner, R. D.; Ogryzlo, E. A. *J. Chem. Phys.* **1984**, *80*, 1.
- (9) Caubet, Ph.; et al. *Chem. Phys. Lett.* **1984**, *108*, 217.
- (10) Chu, A.-L.; et al. *J. Phys. Chem.* **1986**, *90*, 466.
- (11) Viereck, R. A.; et al. *Nature* **1991**, *354*, 48, and references therein.
- (12) (a) Anketell, J.; Nicholls, R. W. *Rep. Phys.* **1970**, *33*, 269.
- (b) Partridge, H.; et al. *J. Chem. Phys.* **1988**, *88*, 3174, and references therein.
- (13) Benesch, W. *J. Chem. Phys.* **1983**, *78*, 2978.
- (14) Boef, J. P.; Kunhardt, E. E. *J. Appl. Phys.* **1986**, *60*, 915. Armenise, I.; et al. *Chem. Phys. Lett.* **1992**, *200*, 597.
- (15) Cerny, D.; et al. *J. Phys. Chem.* **1981**, *85*, 2626.
- (16) Luckovsky, G.; et al. *J. Vac. Sci. Technol.* **1986**, *A4*, 681.
- (17) Weinreb, M. P.; Mannella, G. G. *J. Chem. Phys.* **1969**, *51*, 4973.
- (18) Brennen, W.; Mc Intyre, P. *Chem. Phys. Lett.* **1982**, *90*, 457.
- (19) Kamaratos, E. Unpublished results from (a) York University, CRESS, Toronto Canada, lab. of Prof. H. I. Schiff; communications with Profs. C. H. Dugan, H. O. Pritchard, R. A. Young of CRESS; (b) University of Ioannina, Greece; (c) Université de Lyon I, France, lab. of Prof. R. Bacis.
- (20) Fontijn, A.; et al. *J. Chem. Phys.* **1964**, *40*, 64.
- (21) Piper, L. G.; et al. *J. Chem. Phys.* **1981**, *75*, 2847.
- (22) Meyer, J. A.; et al. *J. Phys. Chem.* **1970**, *74*, 2238.
- (23) Piper, L. G. *J. Chem. Phys.* **1982**, *77*, 2373.
- (24) Thomas, J. M.; Kaufman, F. *J. Chem. Phys.* **1985**, *83*, 2900.
- (25) Heidner, R. F., III; et al. *Chem. Phys. Lett.* **1976**, *37*, 243.
- (26) Sadeghi, N.; Setzer, D. W. *Chem. Phys. Lett.* **1981**, *77*, 304, and references therein.
- (27) (a) Chernysheva, N. V.; Ionikh, Yu. Z. *Opt. Spectrosc. (USSR)* **1979**, *47*, 37. (b) Rotem A.; et al., *J. Chem. Phys.* **1982**, *76*, 2109; and references in a and b.
- (28) Bachmann, R.; et al. *J. Chem. Phys.* **1992**, *96*, 5151, and references therein.
- (29) (a) Piper, L. G. *J. Chem. Phys.* **1989**, *91*, 864. (b) Katayama, D. H.; Dentamaro, A. V. *J. Chem. Phys.* **1992**, *97*, 2820. (c) Scriptor, C.; et al. *Chem. Phys. Lett.* **1993**, *201*, 194, and references therein. (d) Ottinger, Ch.; Vilesov, A. F. *Chem. Phys. Lett.* **1993**, *211*, 175.
- (30) Black, G.; Slanger, T. G. *J. Chem. Phys.* **1981**, *74*, 6517.
- (31) Westenberg, A. A.; et al. *Chem. Phys. Lett.* **1970**, *7*, 597.
- (32) Clark, I. D.; Wayne, R. P. *Chem. Phys. Lett.* **1969**, *3*, 405.
- (33) Schmidt, C.; Schiff, H. I. *Chem. Phys. Lett.* **1973**, *23*, 339.
- (34) Thomas, J. M.; et al. *J. Chem. Phys.* **1987**, *86*, 6885.
- (35) Oldman, R. J.; Broida, H. P. *J. Chem. Phys.* **1969**, *51*, 2254.
- (36) Schiff, H. I. *Ann. Geophys.* **1969**, *25*, 815.
- (37) Schiff, H. I. *Ann. N.Y. Acad. Sci.* **1970**, *171*, 188.
- (38) Because of relocation finally to Greece and being involved in higher energy transfer studies [*Chem. Rev.* **1984**, *84*, 561; *Physica* **1985**, *132C*, 260], I learned about the recent investigations too much later.
- (39) Kamaratos, E. Submitted for publication.
- (40) APS Study Group, *Rev. Mod. Phys.* **1987**, *59*, S1.